NN-based inference of the optical depth to reionization τ from Planck maps¹ Kevin Wolz (SISSA, INFN Trieste), Nicoletta Krachmalnicoff (SISSA, INFN, IFPU Trieste),

Background

- The optical depth to reionization, τ , is the least constrained ACDM parameter. To date, its most precise value stems from large-scale EE cross power spectra (XCL) from Planck's highfrequency instrument (HFI).
- maps contain hard-to-model non-Gaussian Planck HFI systematic residuals. Robust XCL methods, using simulations to build an empirical likelihood, yield $\tau = 0.0566^{+0.053}_{-0.062}$ (68% CL).²
- We present the first likelihood-free τ inference from polarized *Planck* HFI maps fully based on neural networks (NNs).

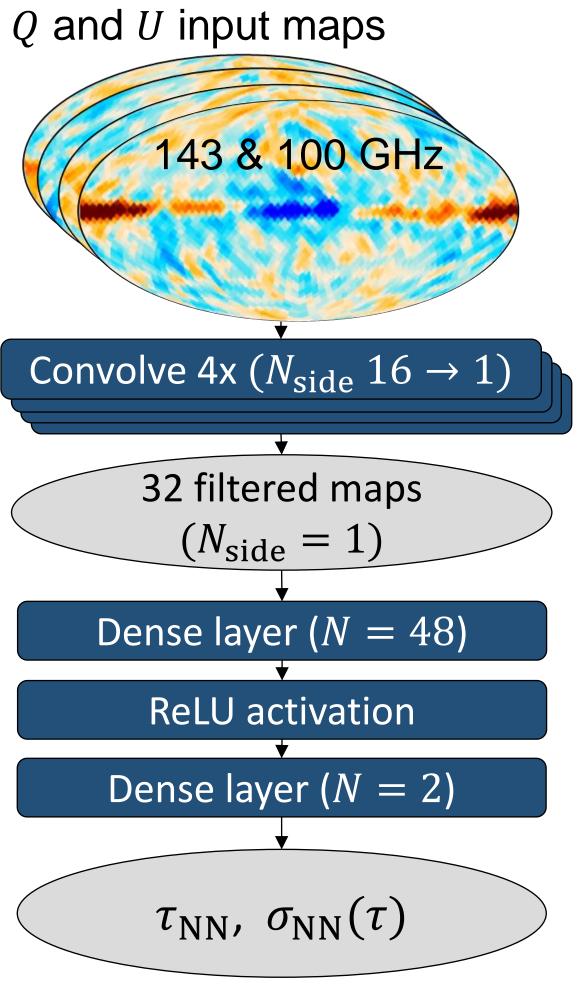
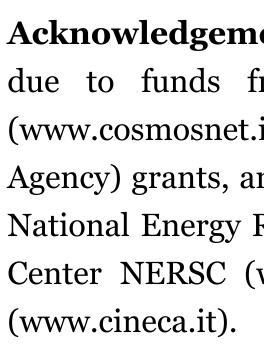


Fig. 1. Schematic of the convolutional dense neural net used in this work.

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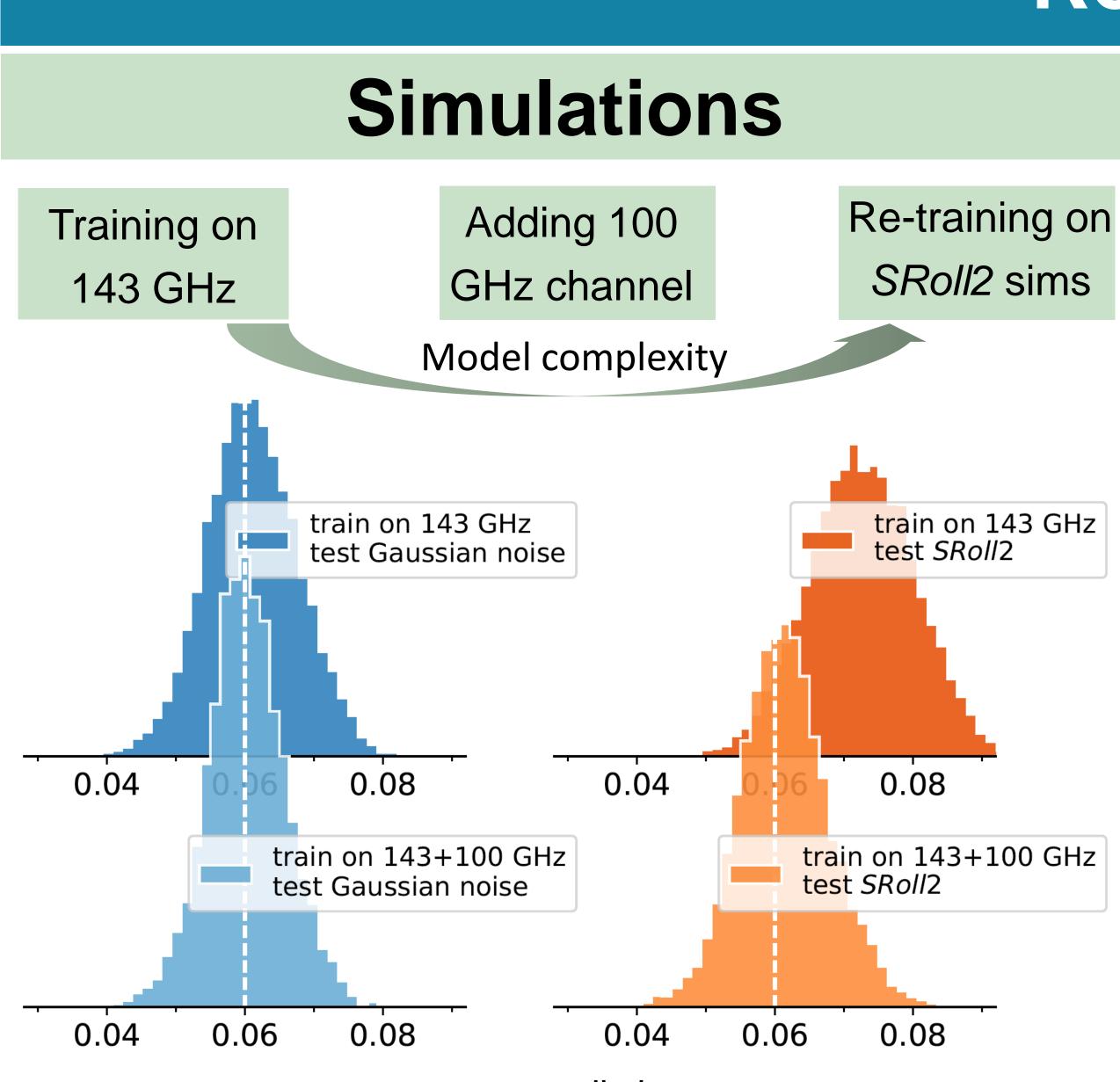
Method: CNNs

- Convolutional neural nets (CNNs) do not need analytical modeling and can be trained on multi-channel sims.
- We train, test, and validate NN models on Gaussian or *Planck* SRoll2³ sky sims (CMB, noise, systematics).
- Infer value of τ directly from Stokes Qand U maps at $\sim 4^{\circ}$ pixel resolution, convolving first-neighbor pixels on the sphere using the *NNhealpix* code.⁴



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 $[\]tau_{NN}$ prediction

Fig. 2. Predictions from 10⁴ simulations with input τ = 0.06, of CMB with Gaussian noise (left panels) or CMB with SRoll2 noise + systematics (right panels). CNN models are trained on CMB + Gaussian noise on one (top) or two frequency channels (bottom).

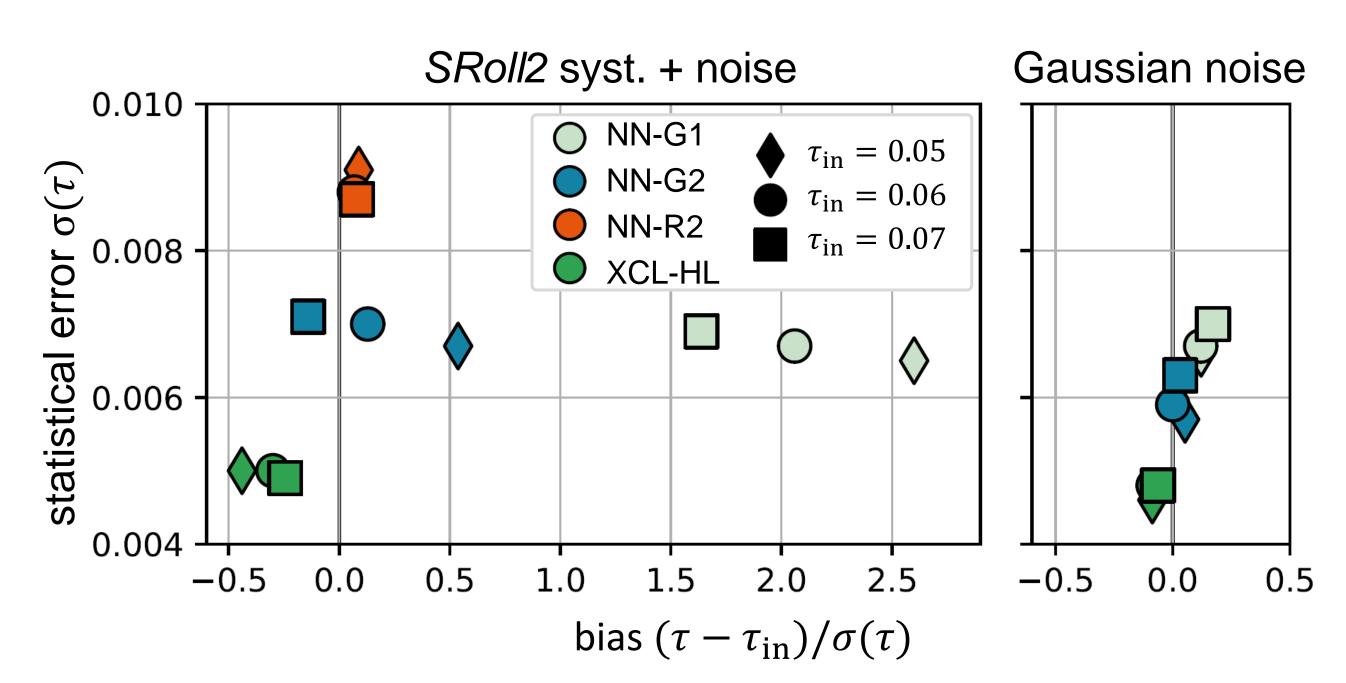
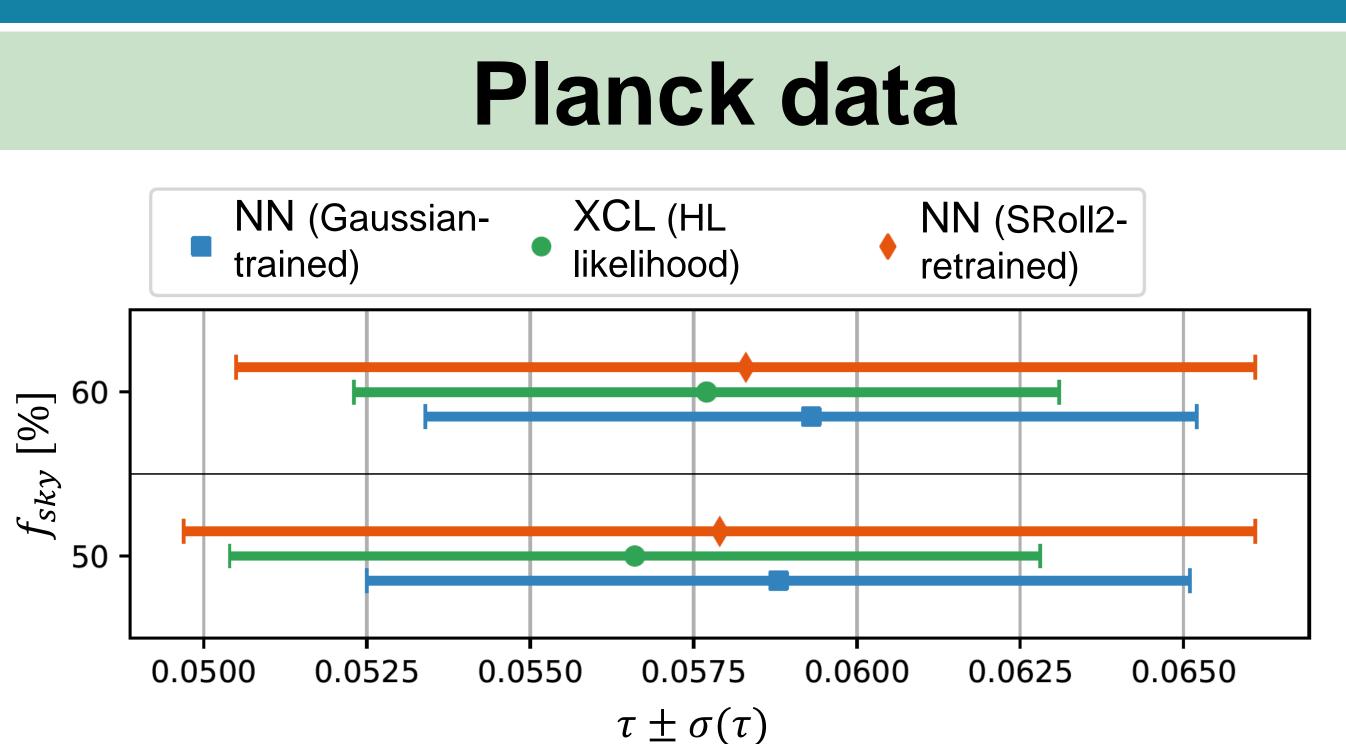


Fig. 3. Prediction bias and statistical error from sims as in Fig. 2, with different input τ values. CNN models are trained on 1 or 2 channels of Gaussian noise (teal, blue) or retrained on SRoll2 sims (orange). For comparison, we show results from the Hamimeche-Lewis (HL) XCL likelihood (green).

Results



Conclusions

- with a ~ 30% larger uncertainty.¹
- NNs without need of explicit modeling.

- ² Luca Pagano et al. 2020, A&A, doi:201936630
- ³ Jean-Marc Delouis et al. 2019, A&A, doi:201834882



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Fig. 4. Results for τ from Planck SRoll2 data, using the standard 50% sky mask or a 60% mask for comparison. Considered are the Gaussian NN, the retrained NN and the HL likelihood.

• On Planck data, we obtain $\tau_{NN} = 0.0579 \pm 0.0082$ (68% CL), compatible with current XCL results but

effectively combine information from two channels, reducing impact of noise and systematics

While this work does not improve $\sigma(\tau)$, it is the first robust NN-based inference on real CMB data.

Promising tool for complementary analysis of nearfuture CMB experiments, *e.g.*, B-mode searches.

References

Kevin Wolz et al. 2023, submitted to A&A, arxiv:2301.09634 ⁴ Nicoletta Krachmalnicoff et al. 2019, A&A, doi:201935211